Multi-scale data assimilation – forecasting high-latitude scintillation

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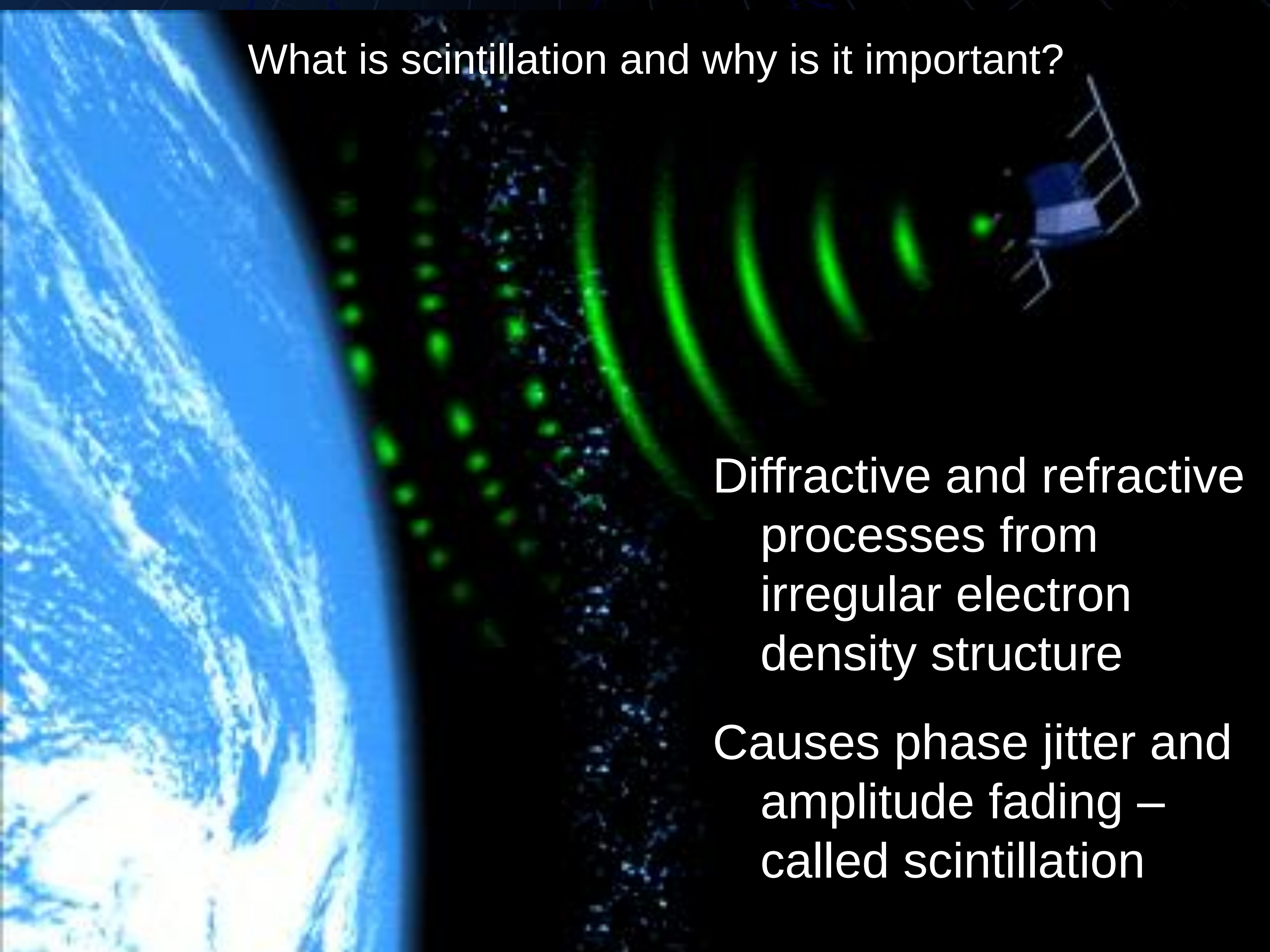
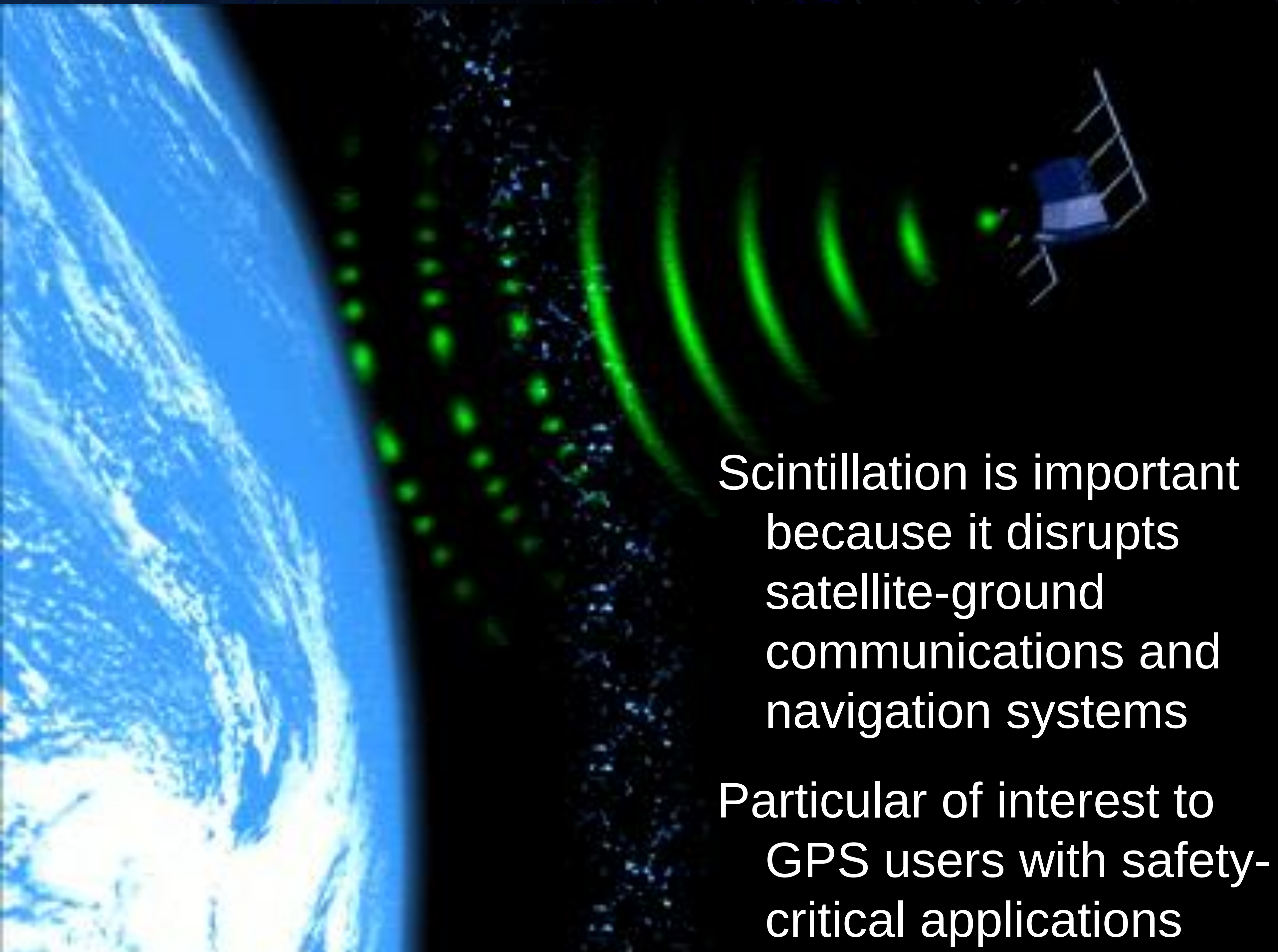
What is scintillation and why is it important?



- Diffractive and refractive processes from irregular electron density structure
- Causes phase jitter and amplitude fading – called scintillation


Scintillation is important because it disrupts satellite-ground communications and navigation systems
Particular of interest to GPS users with safety-critical applications

Scintillation

Scintillation varies widely in significance – some users will see no effect whereas other will suffer complete signal loss

Two indices used to quantify the effects:

Sigma phi quantifies **phase** scintillation

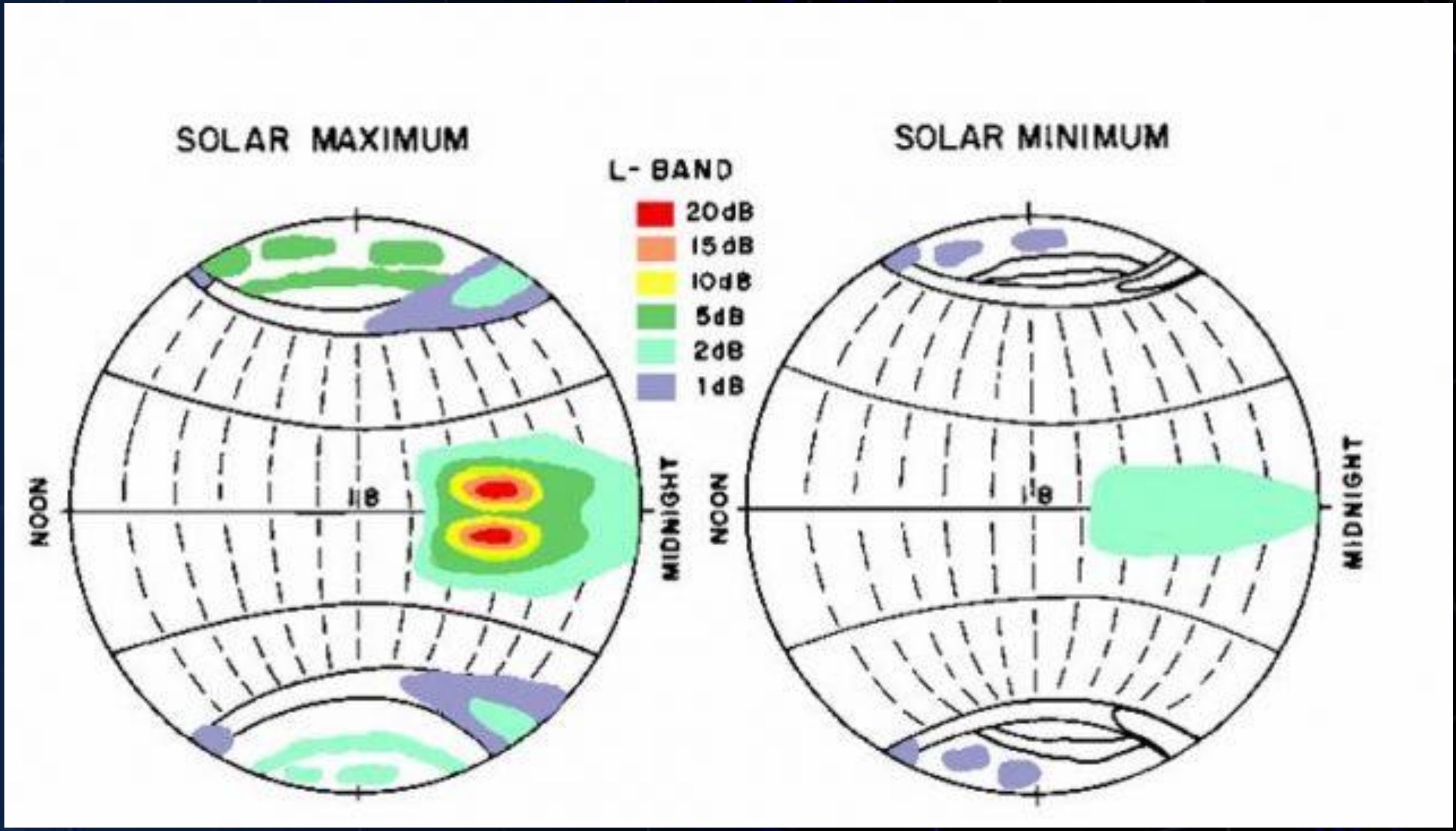
Phase scintillation is more common at **high latitudes**

S4 quantifies **amplitude** scintillation

Amplitude scintillation is more common at **equatorial latitudes**

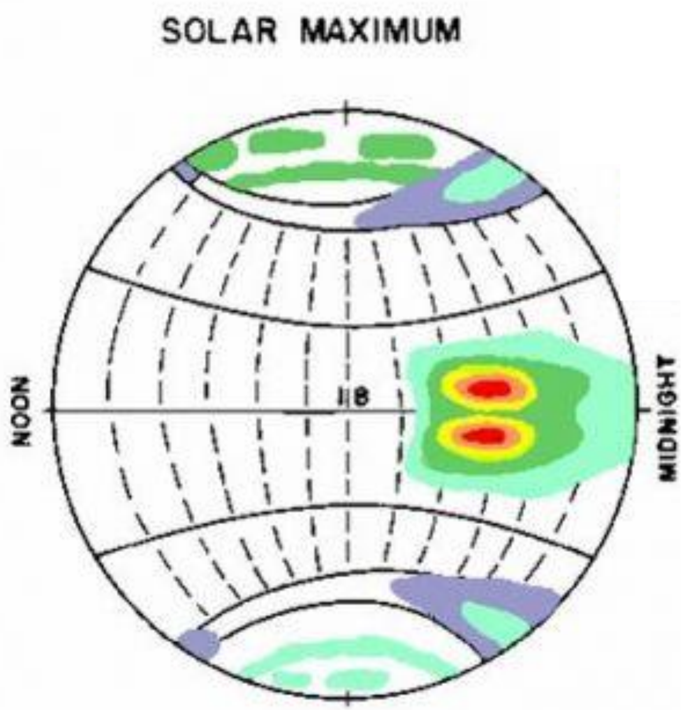
Scintillation is more severe at lower operating frequencies

Global Temporal and Solar Cycle View



Summary picture of scintillation activity at GPS frequencies (after Basu et al)

GPS Scintillation equipment deployed as of 2012



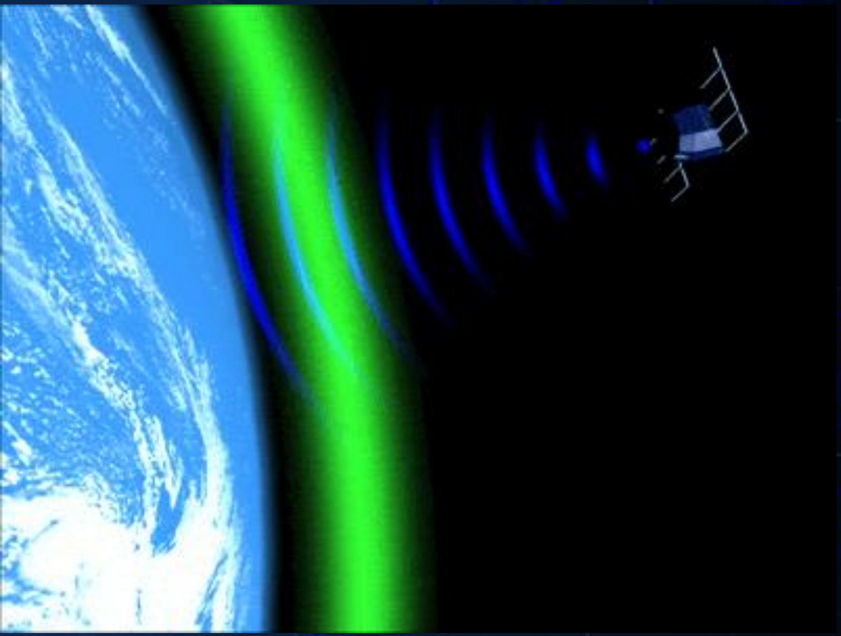
20+ receivers

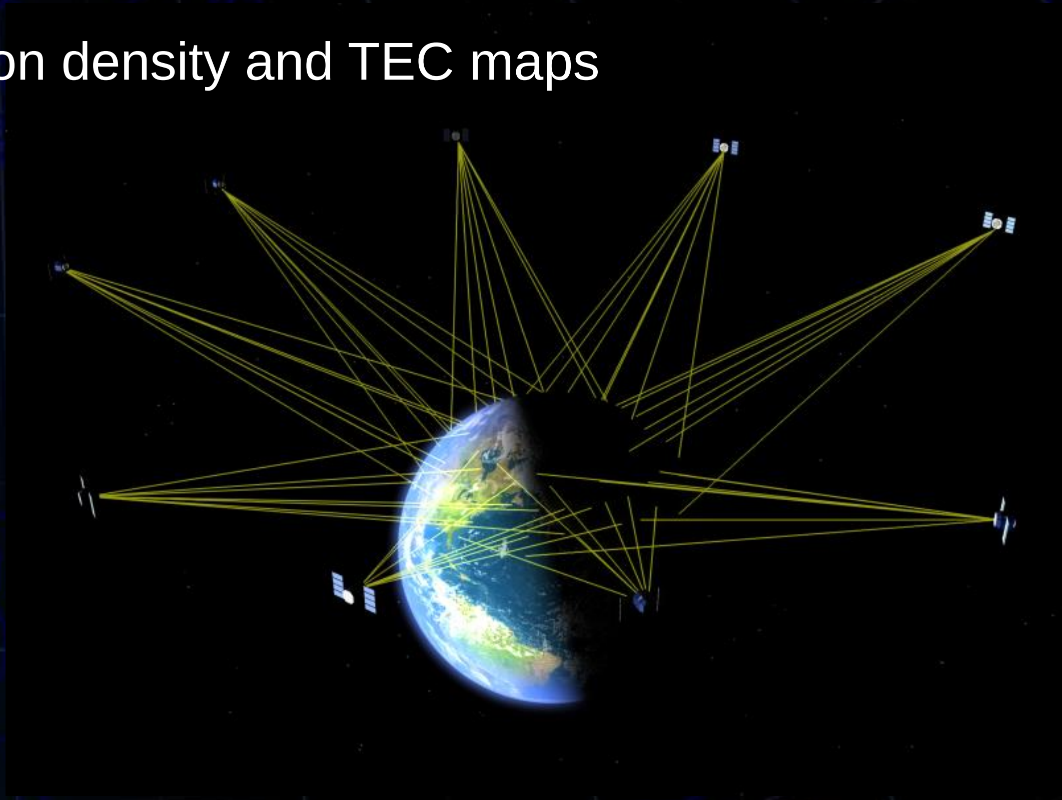
70+ receivers

10+ receivers

Ionospheric Imaging / Data Assimilation

Production of electron density and TEC maps





Geodetic GPS data inputs, IMF, Kp, Kalman filter

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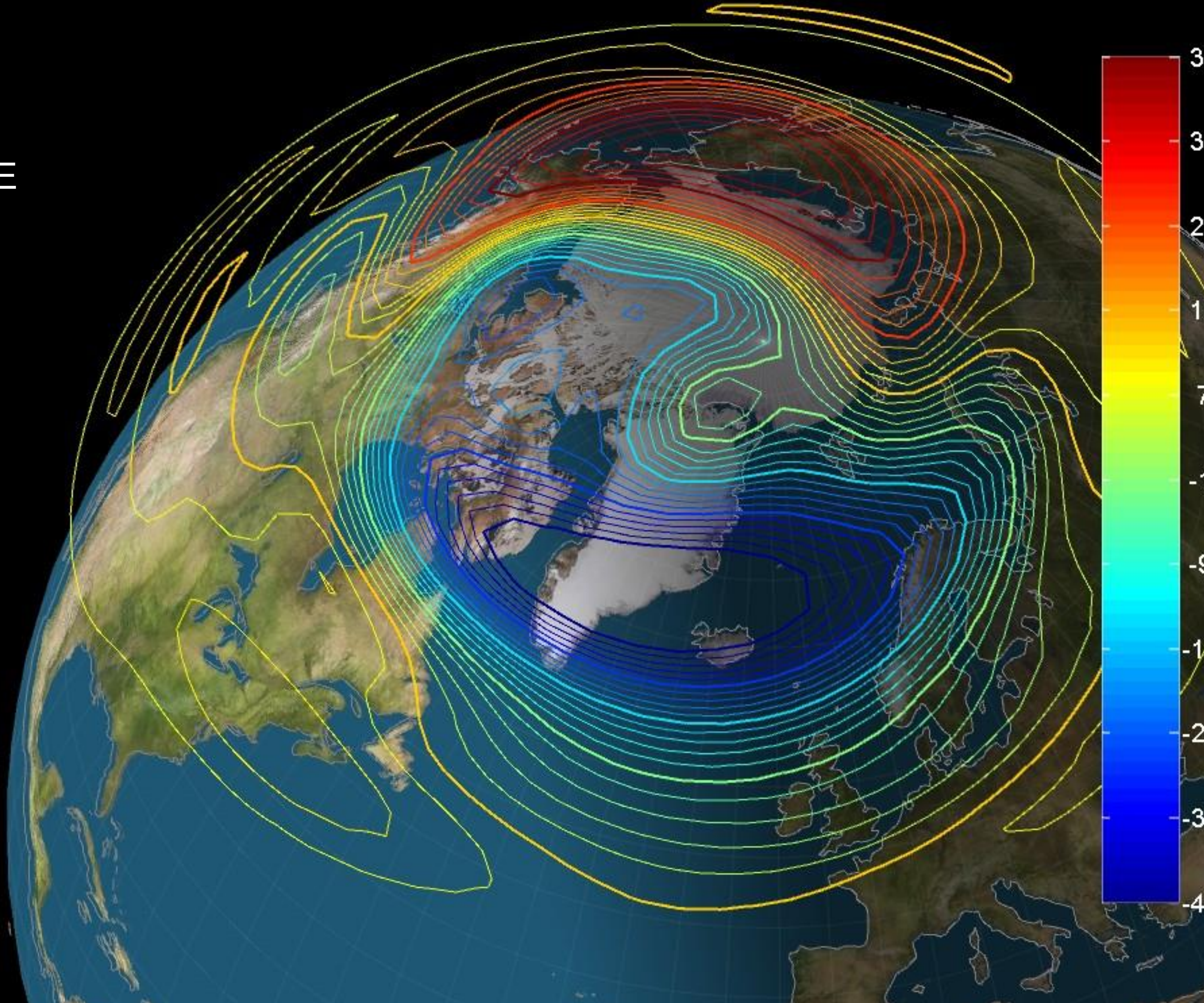
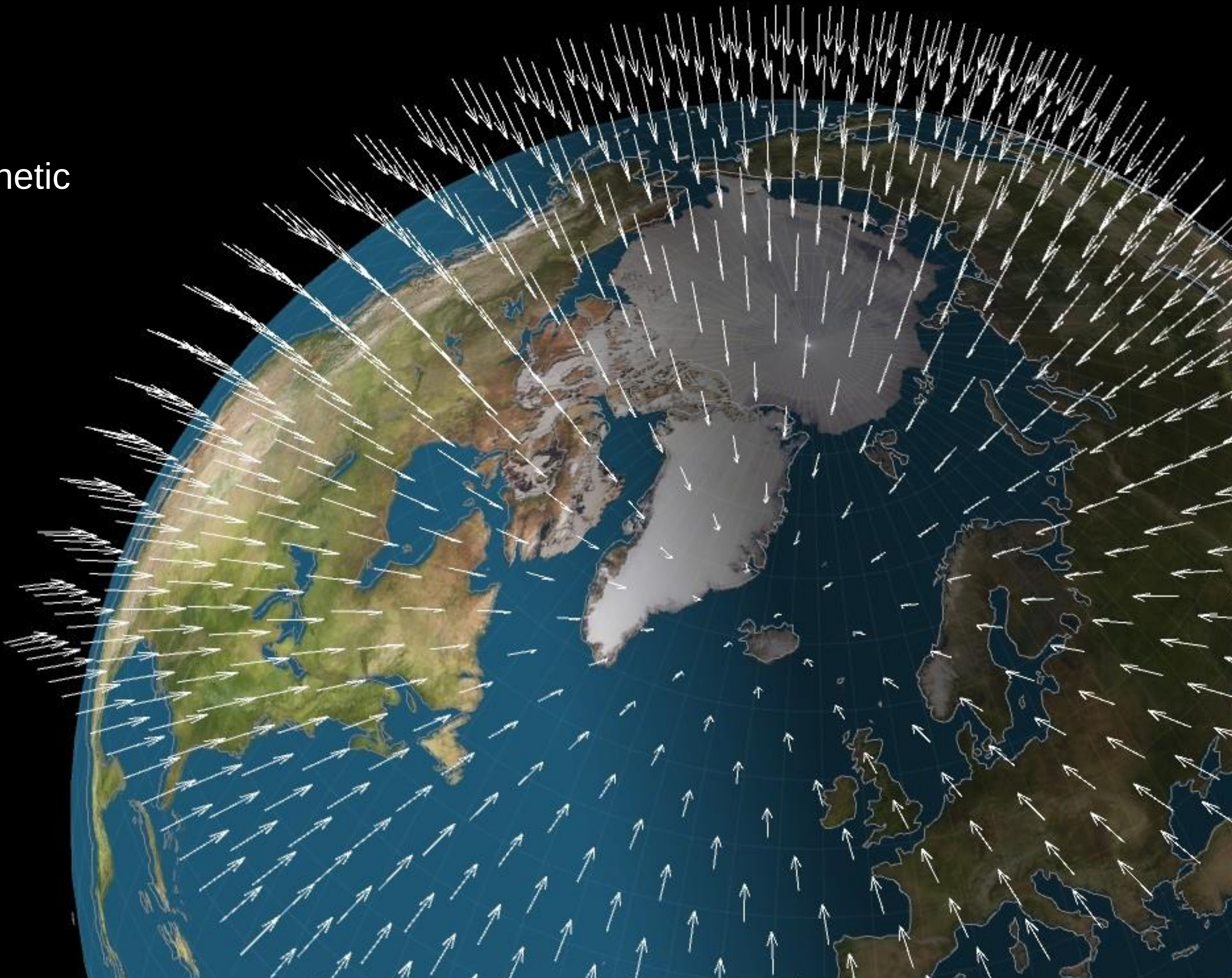
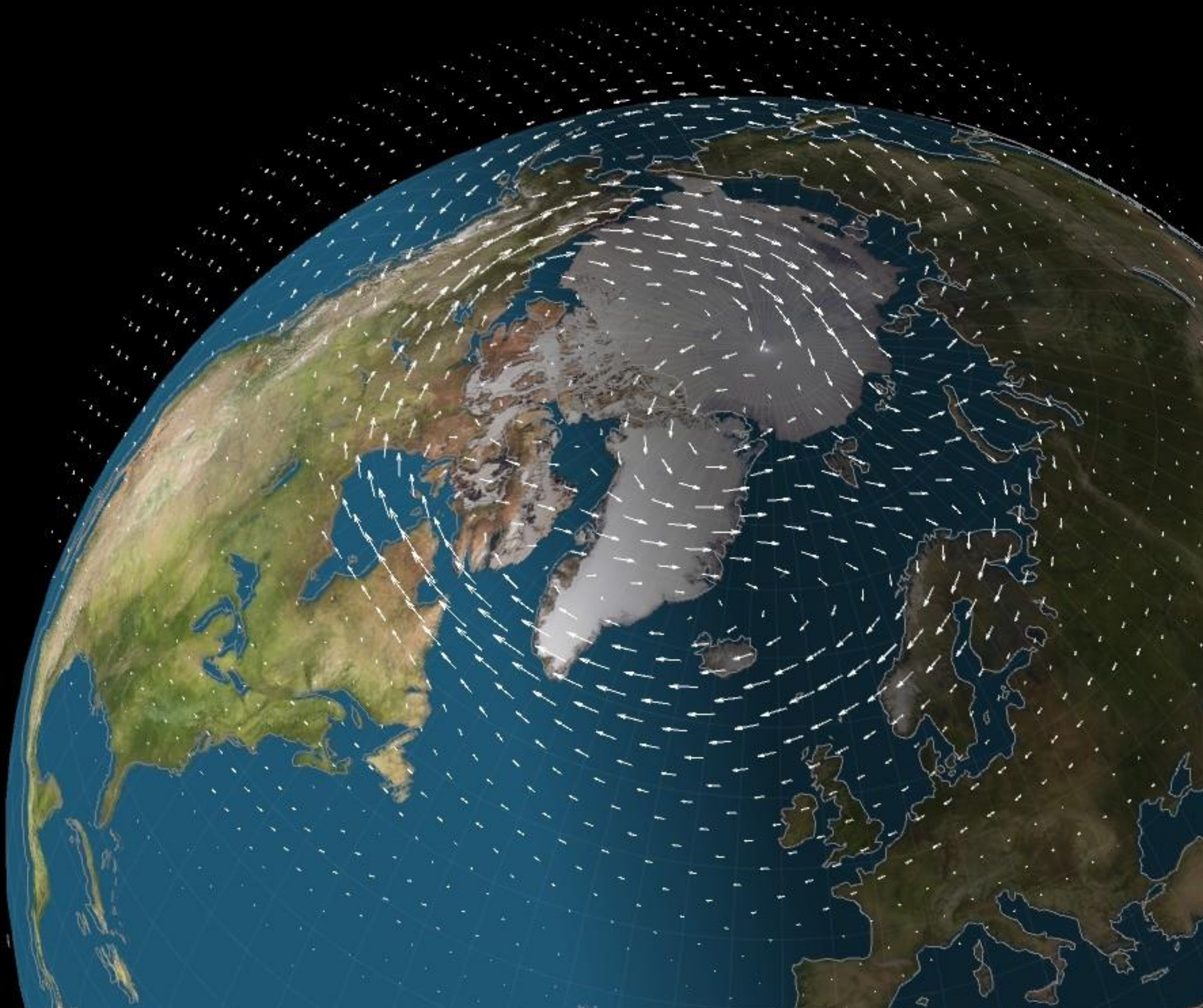
Weimer E field



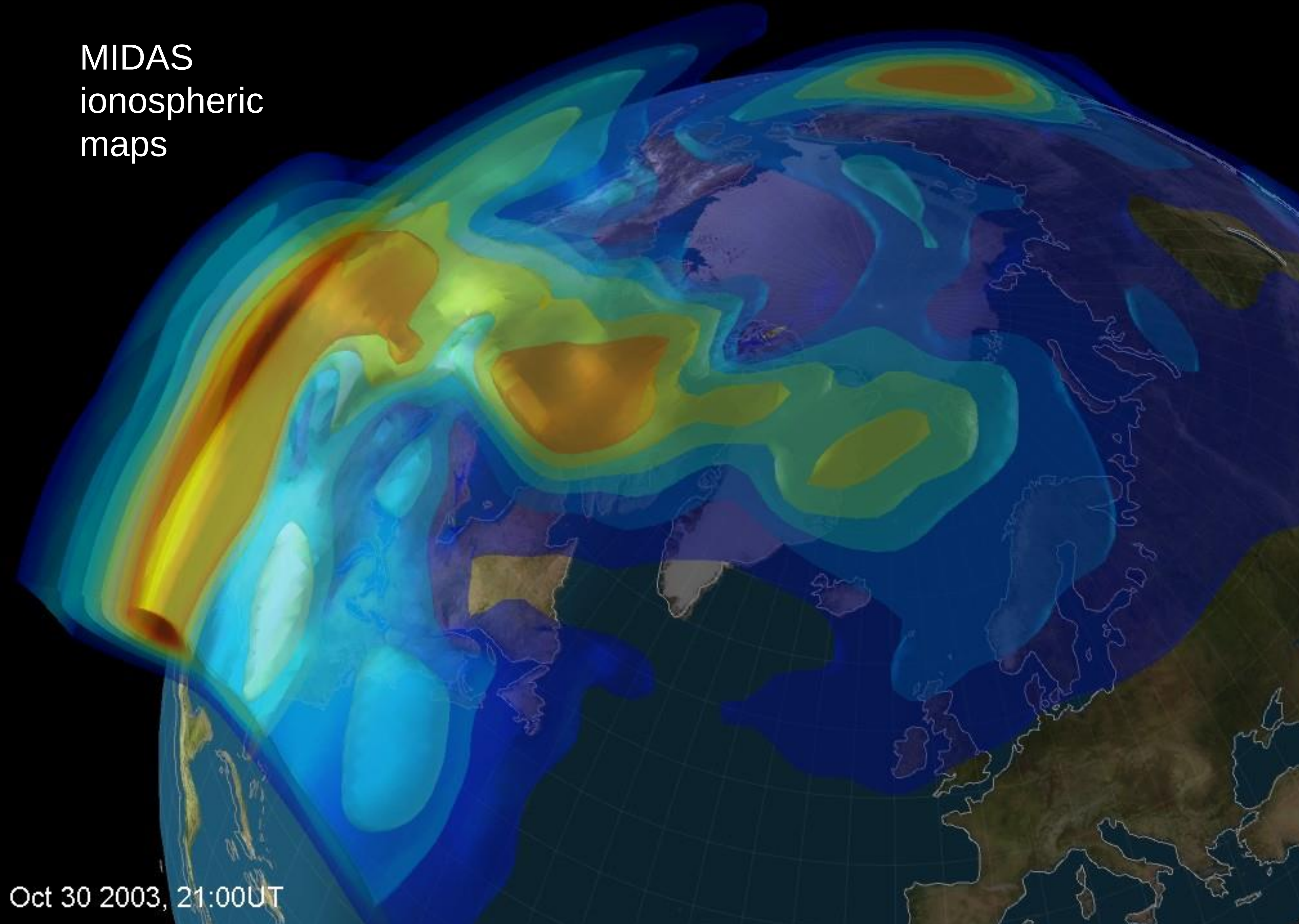
Magnetic
field



Velocity



MIDAS
ionospheric
maps



Relate to Scintillation

Example showing scintillation on the edges of polar cap patches

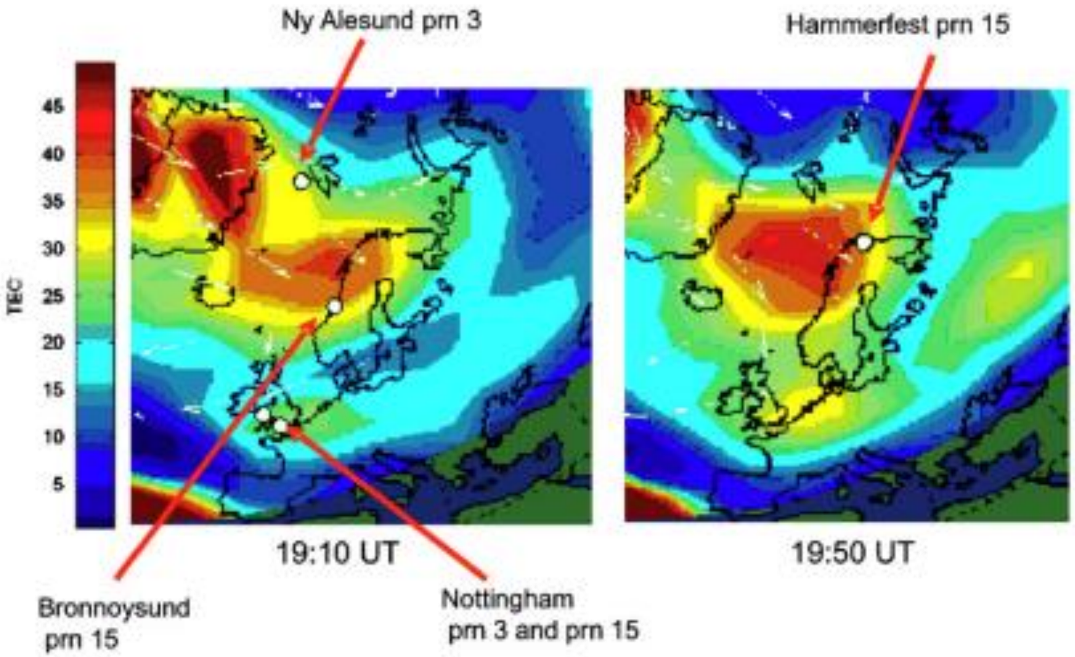
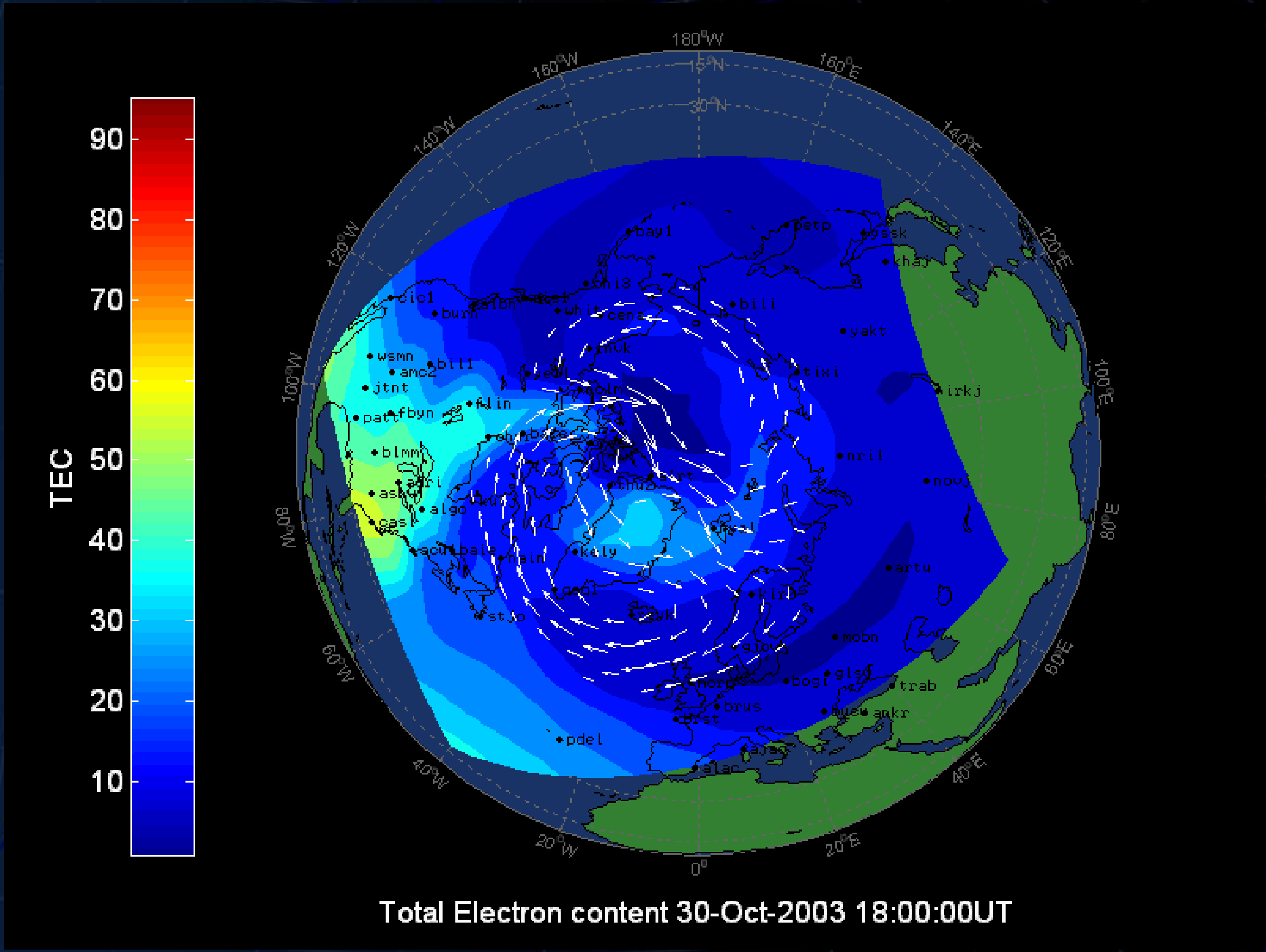


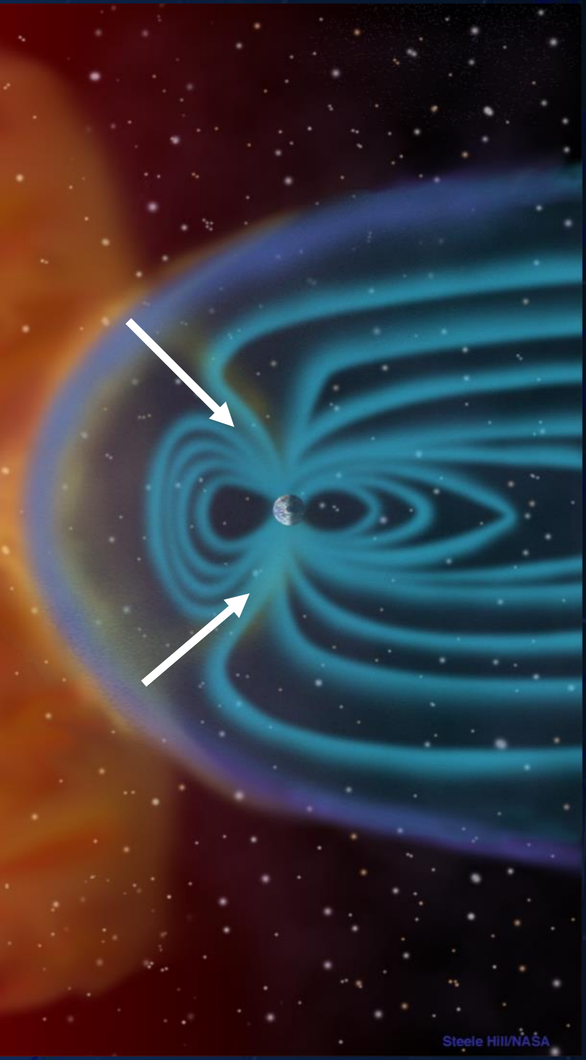
Fig. 4. Equivalent vertical TEC (TECU) snapshots by MIDAS for 30 October at 21:40 and 22:25 UT (top), and 20 November at 19:10 and 19:50 UT (bottom). σ_Φ maxima for selected PRNs as recorded from the GISTM chain are superimposed.

G. De Franceschi et al. / Journal of Atmospheric and Solar-Terrestrial Physics

Possible approach to forecasting high-latitude scintillation

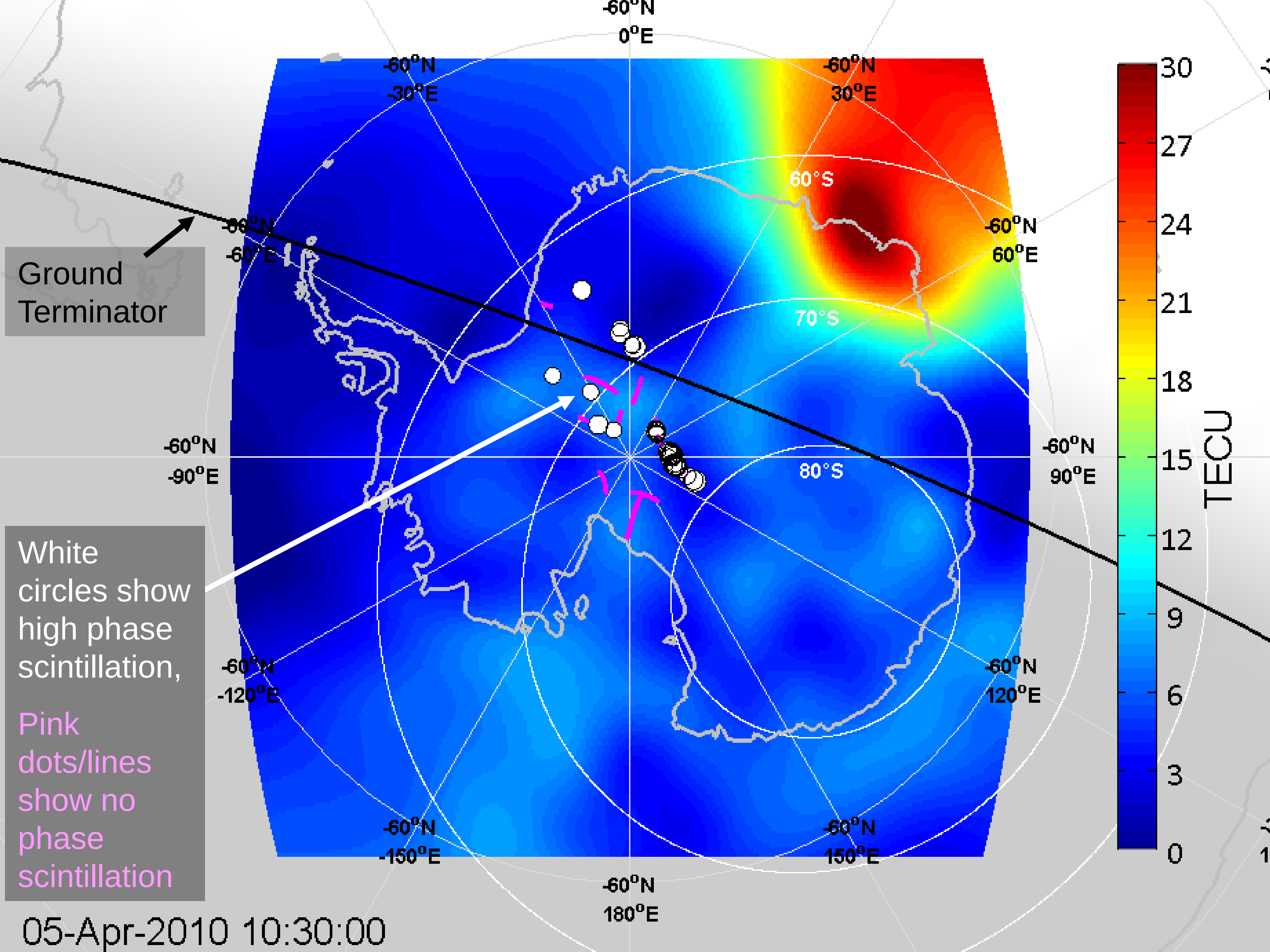


High-latitude scintillation



However, it is not always so simple

Second example:
Scintillation – **dayside** in the cusp



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-60°N
-120°E
-60°N
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-60°N
-150°E
-60°N
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-60°N
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60°S
70°S
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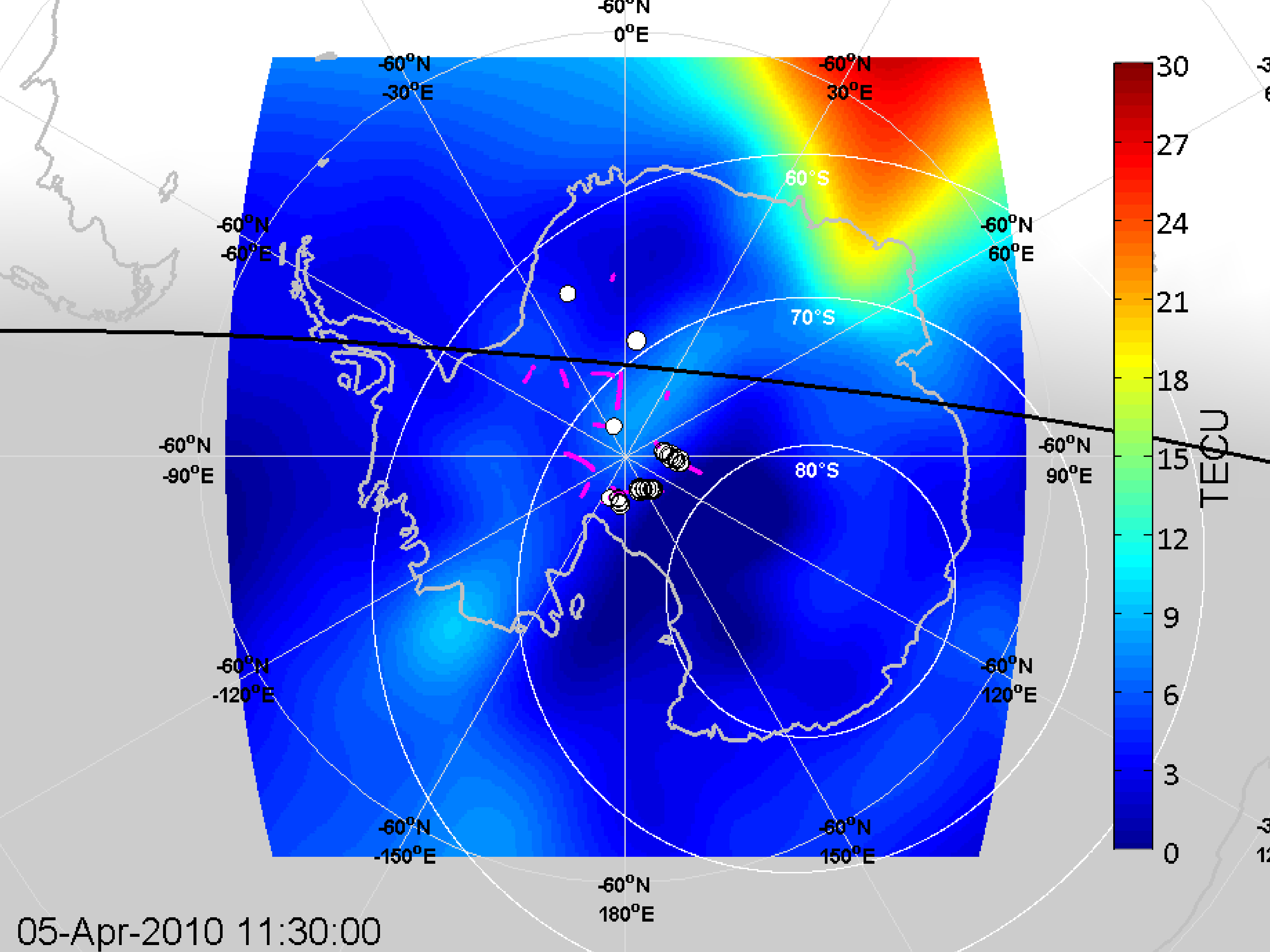
TECU
Ground Terminator
White circles show high phase scintillation,
Pink dots/lines show no phase scintillation
05-Apr-2010 10:30:00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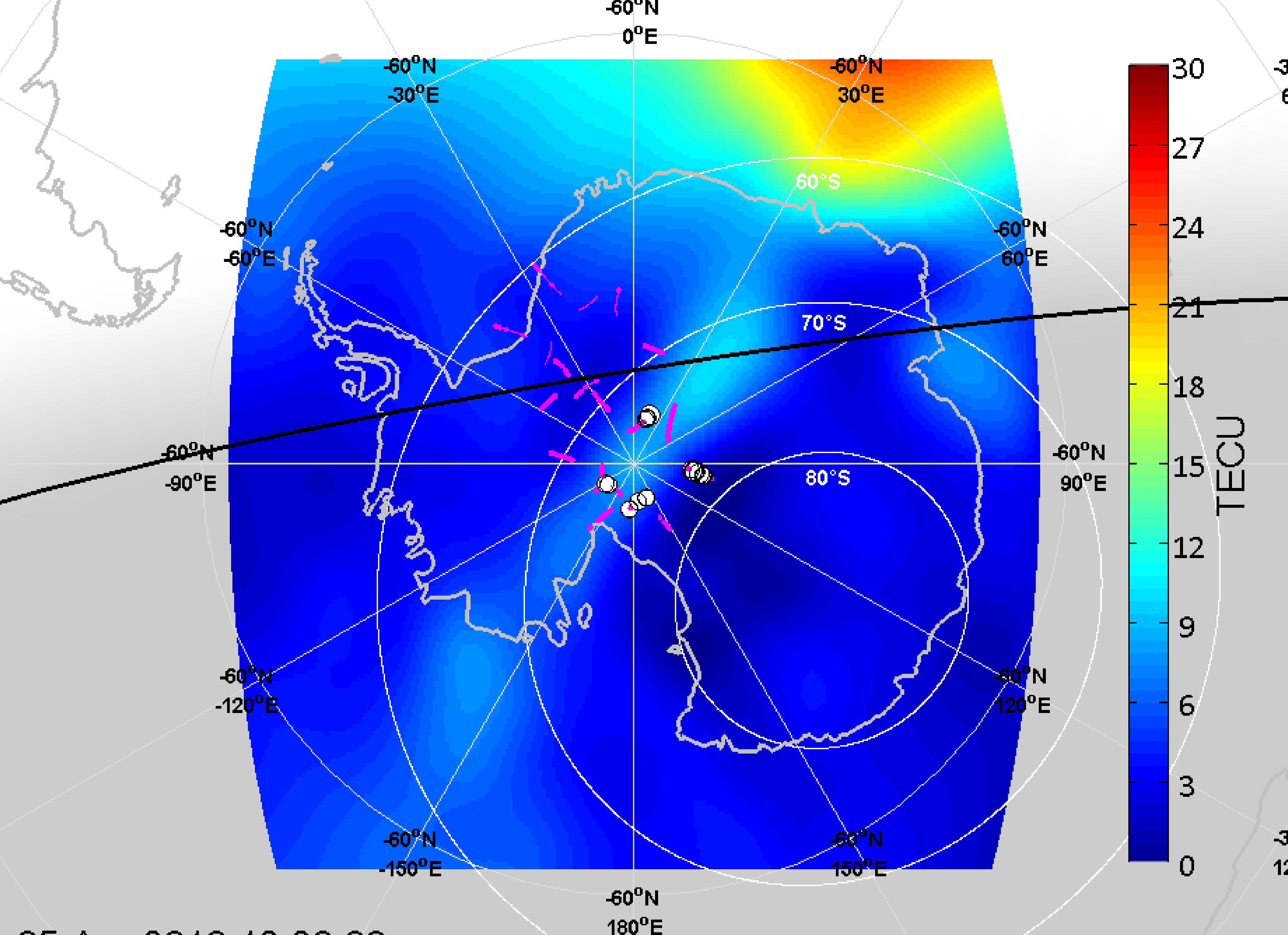
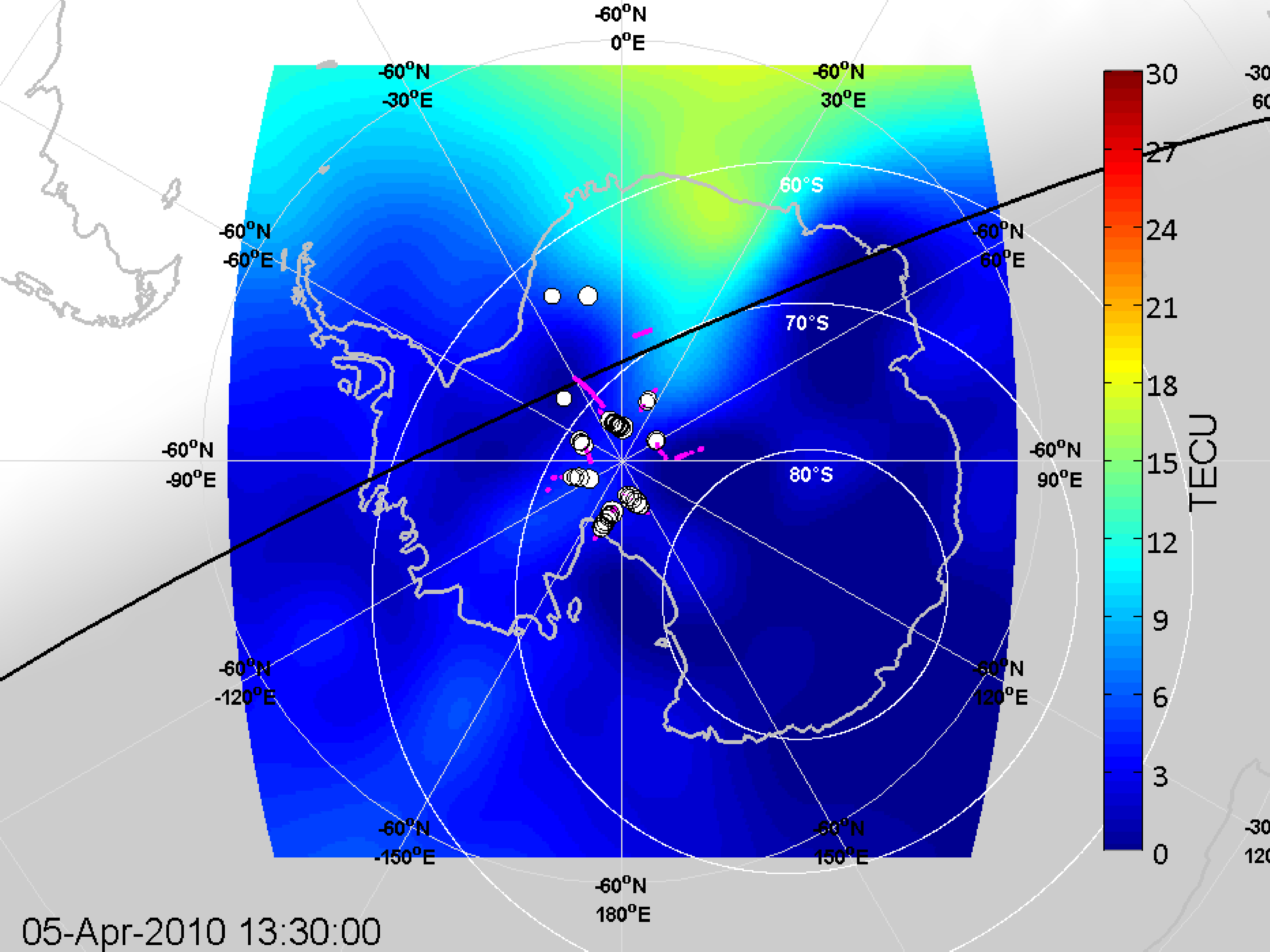


05-Apr-2010 12:30:00



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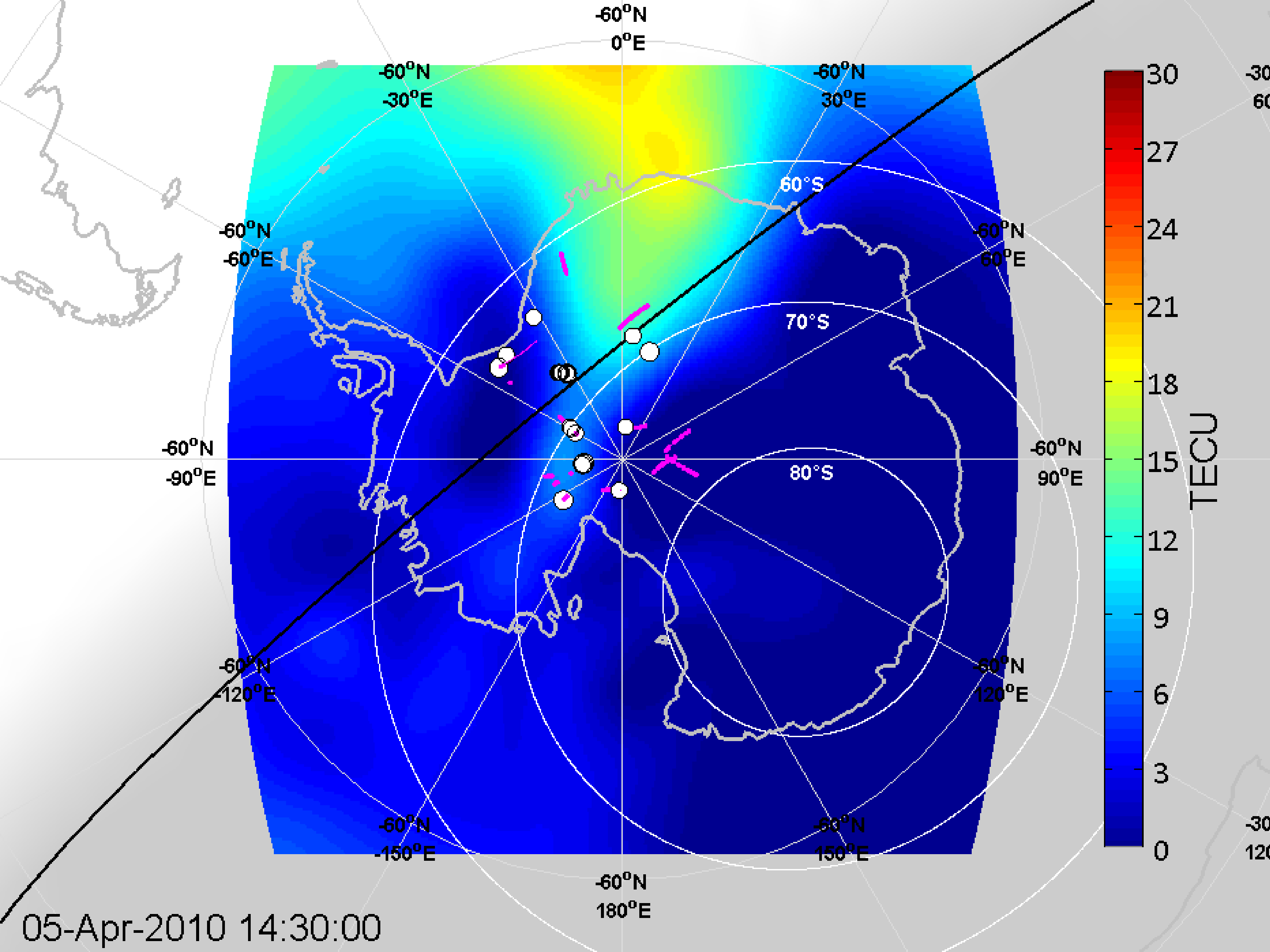

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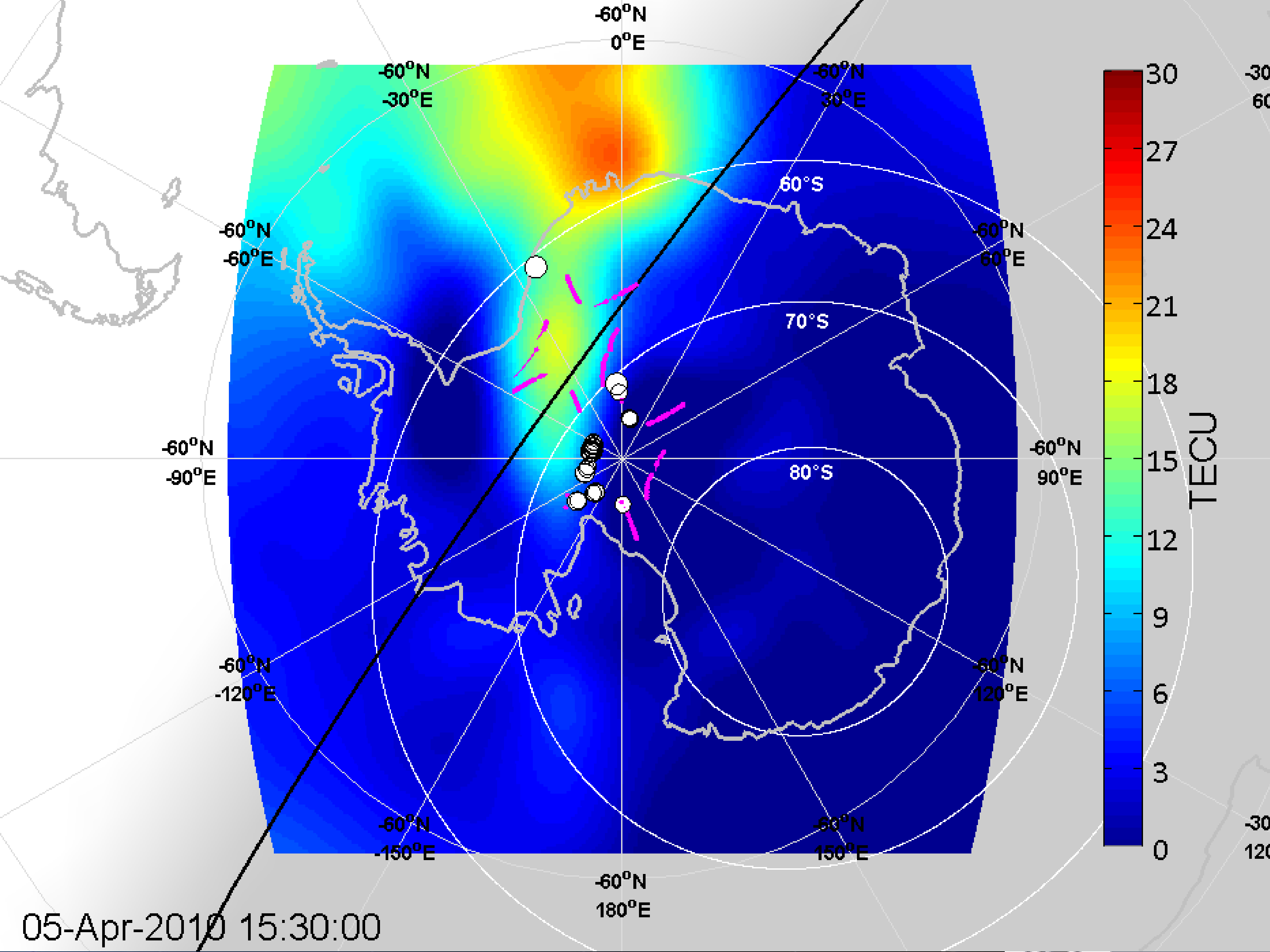

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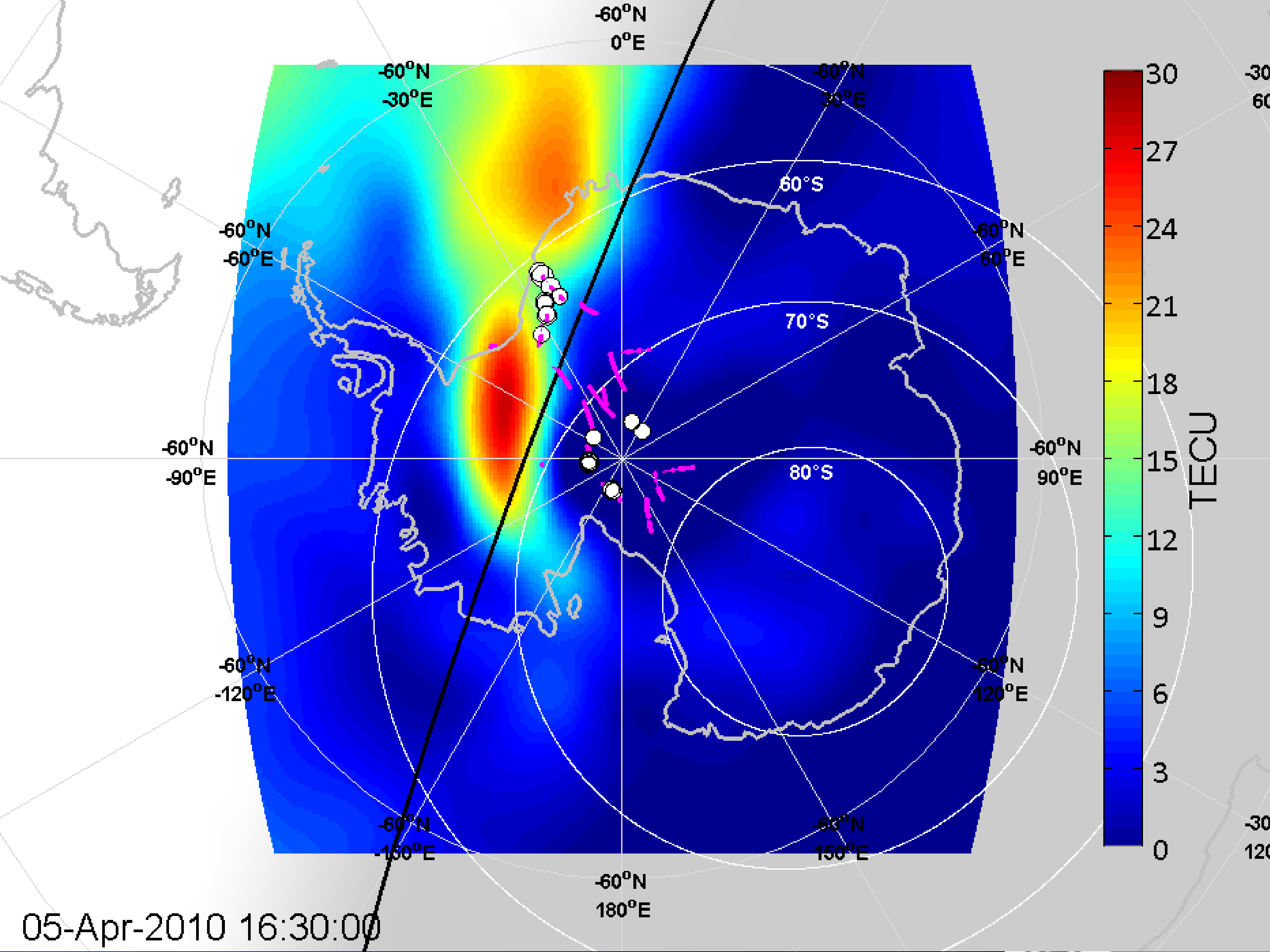

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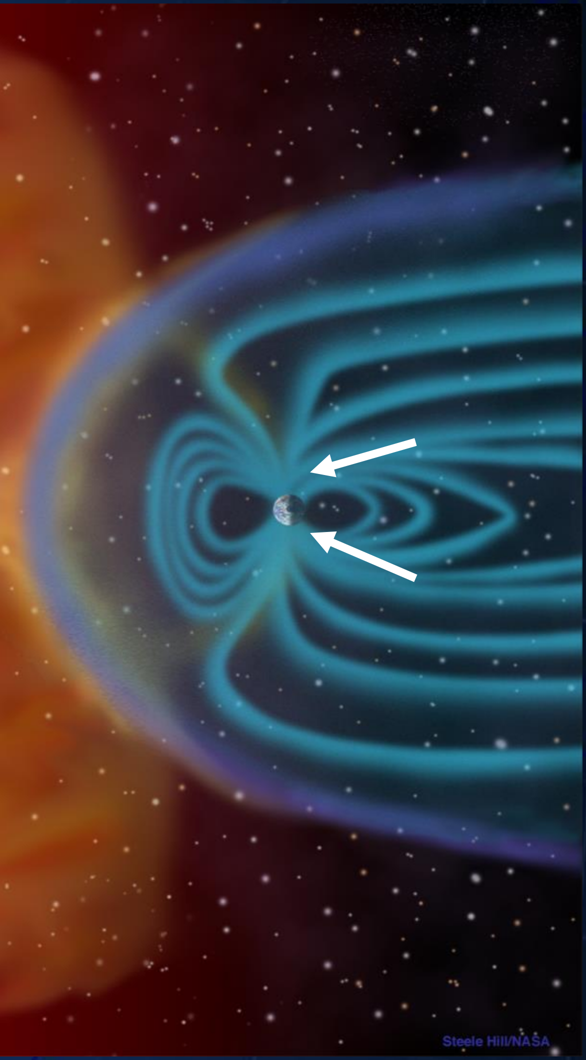
TECU
05-Apr-2010 16:30:00

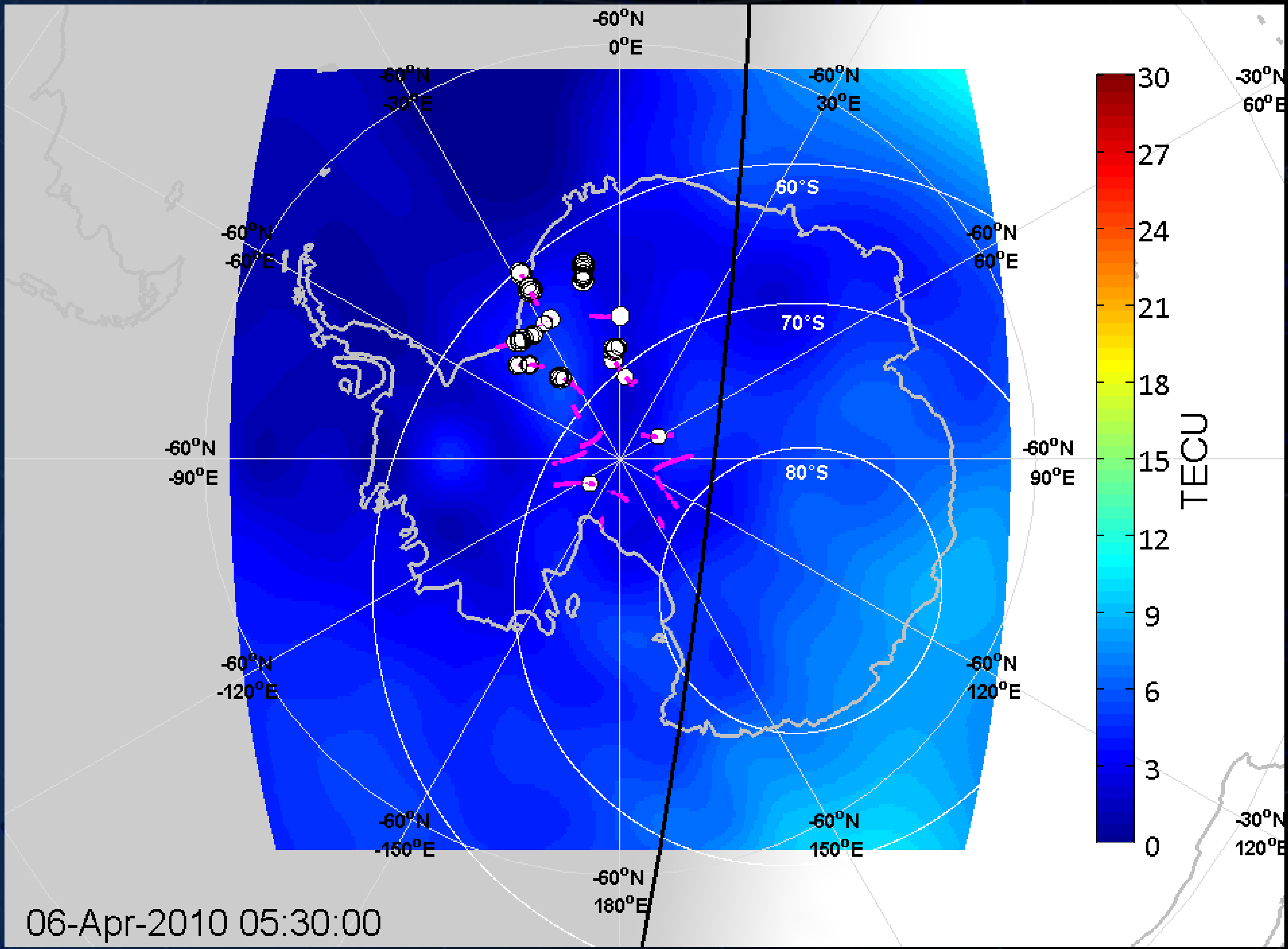
High-latitude scintillation



Third example:

Scintillation – **nightside in the auroral oval**

Nightside on 6 April 2010



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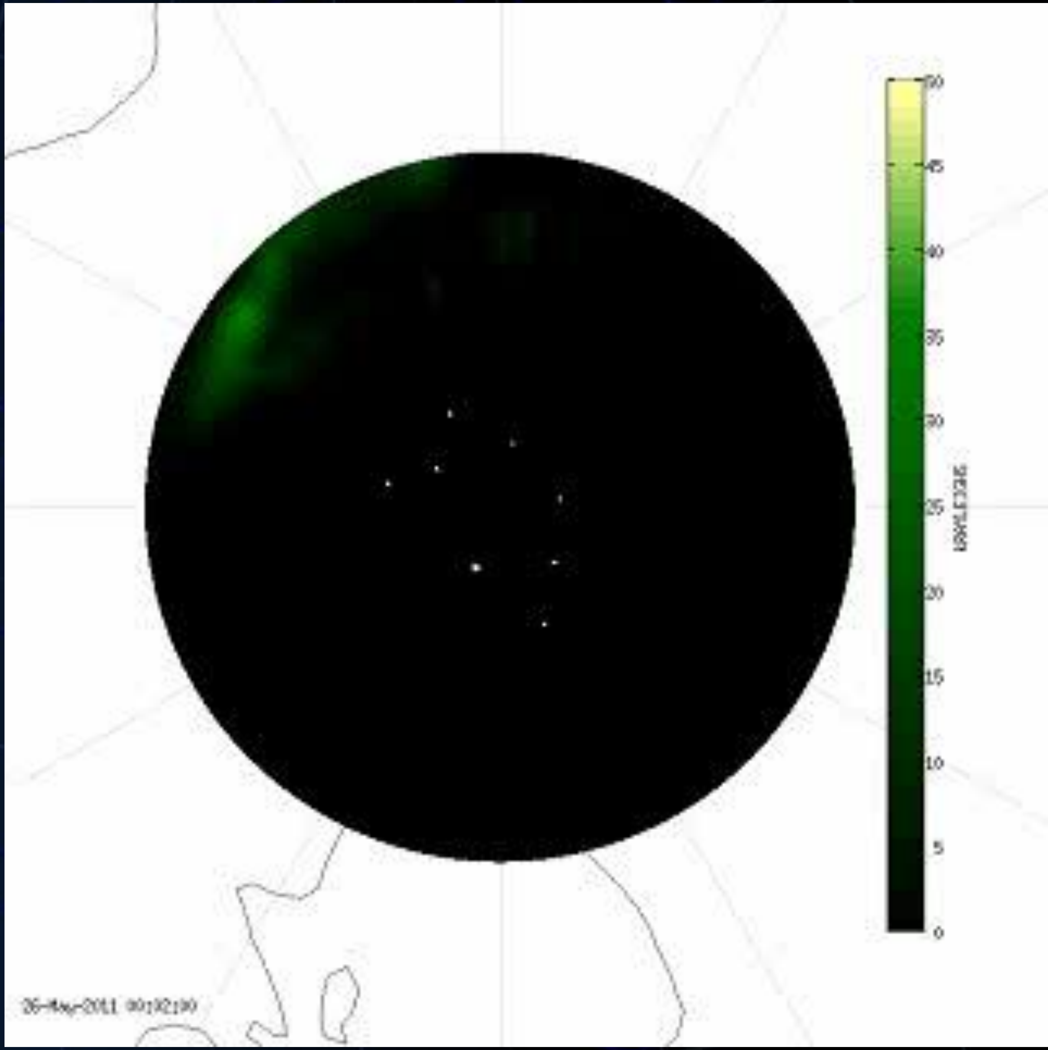
South Pole Station, Field of view above 85 S

26 May 2011

0-2 UT



Acknowledgements: Auroral movie in collaboration with Yusuke Ebihara, Kyoto University, Gary Bust, ASTRA and Al Weatherwax, Siena College.

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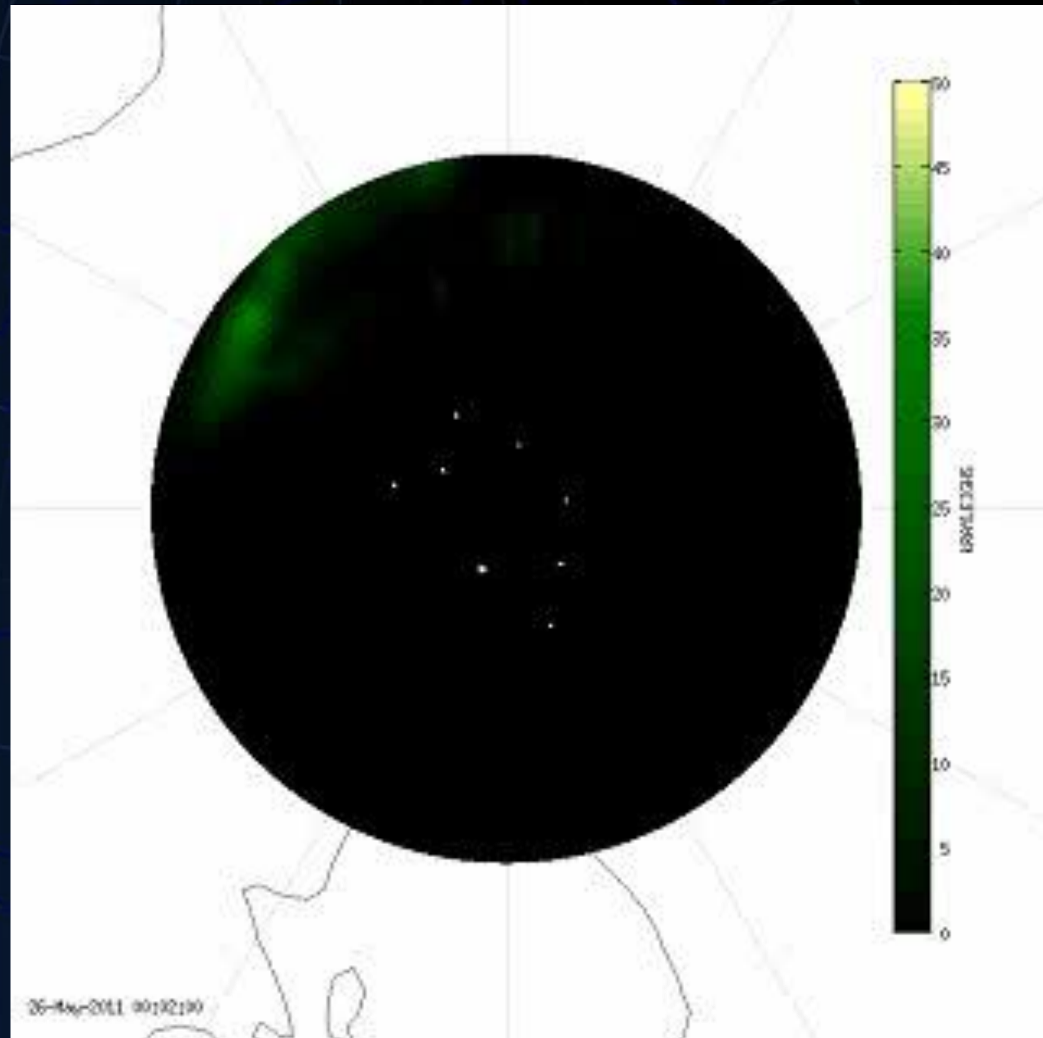
South Pole Station, Field of view above 85 S

26 May 2011

0-2 UT



Difficult to predict which comms/nav link will have a problem

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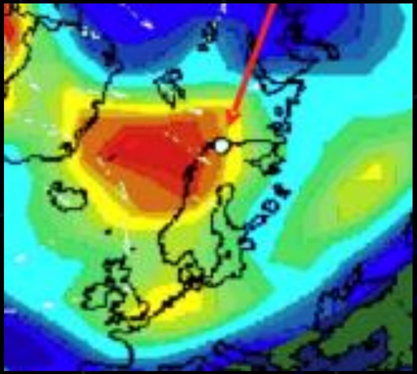
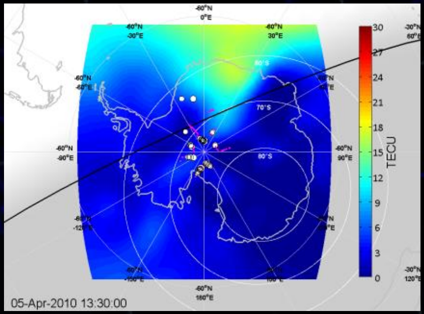
Summary

Three types of high latitude scintillation:

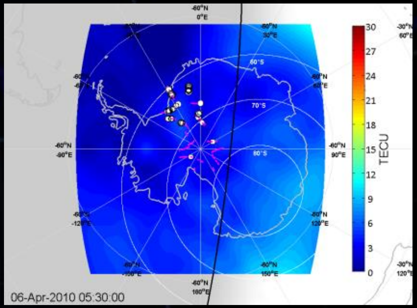
Convecting patches



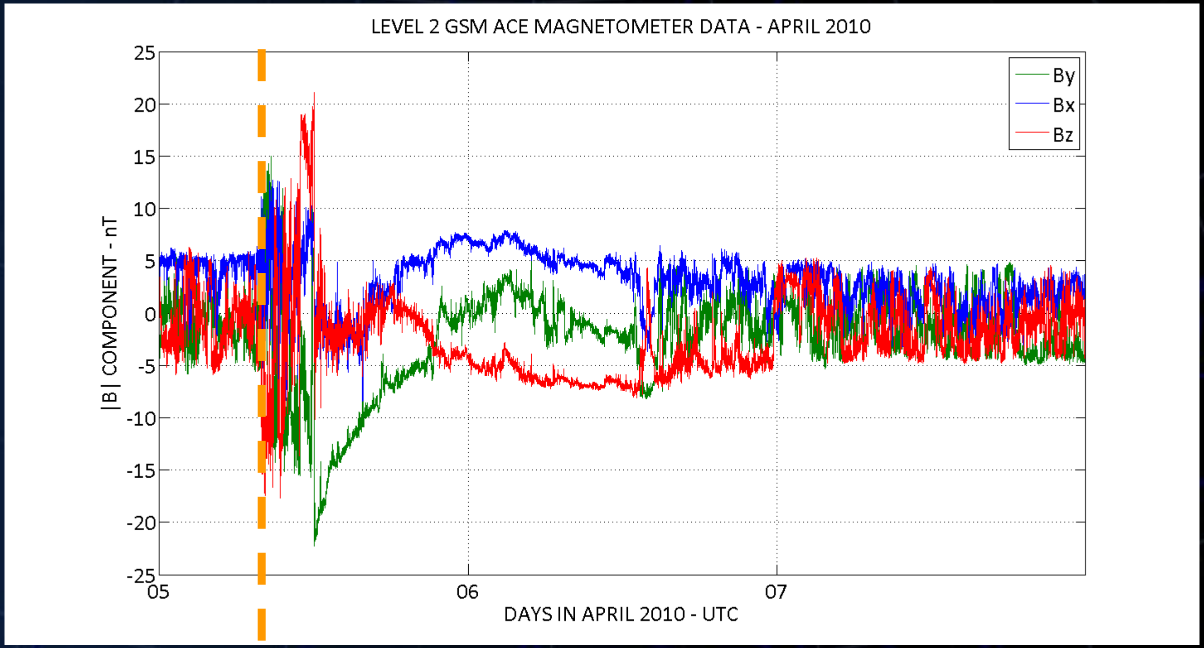
Dayside cusp

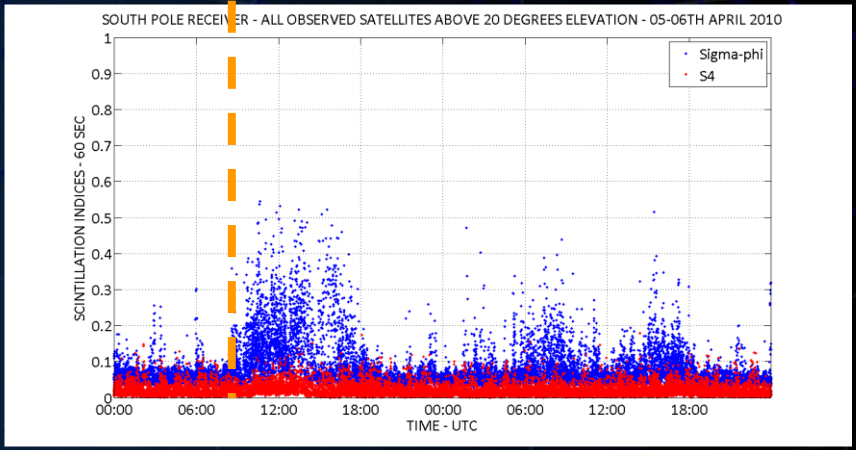


Nightside auroral oval



Scintillation Alert



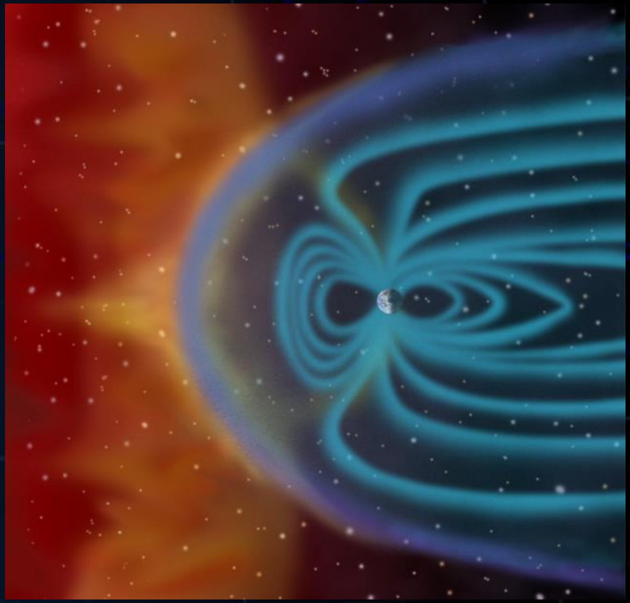


Conclusions

To short-term forecast there are two very different challenges

Patch scintillation – can use ionospheric model to convect

Precipitation scintillation – very difficult to forecast details – magnetosphere probably too late – need to know upstream from ACE and use full data- driven coupled models of Sun-Earth system



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